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# European Transport Solver/Nraining 

P.HEXNHY, Y:BASIUKNIT:ANIEL

## Outline

I. ETS, from the physics point of view -Description
-V\&V
II. ETS, KEPLER workflow
-Description
-How to configure a run from Kepler (as a developer)
-How to add your actor
III. Visualization and Post-treatment
-Matlab
-Python
IV. Practice

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## I. ETS from physics point of view

## ETS, SOLVER (1/4)

$\psi=-\int B d S$

## ITM convention

$\psi$ and $\partial \psi$ quantities have the sign of $-I_{p}$
$q$ has the sign of $-I_{p} *+B_{0}$
$B_{\phi}$ quantities have the sign of $+B_{0}$


## ETS, SOLVER (2/4)

All the ETS transport equations are developed in the following form
where \(\mathrm{F}=\left(\begin{array}{c}\psi <br>
\mathrm{Pe} <br>
\mathrm{Pi} <br>
ne <br>

vtor\end{array}\right\} \quad\) and $\quad \mathrm{X}=\frac{\rho}{\rho_{\max }} \quad$| A, B, C, D are matrices |
| :--- |

The radial grid is uniformed and normalized
The solver takes into account non diagonal terms coupling Pe, Pi and ne Adding a new equation is easy (nbeq=nbeq +1 , needs to fill the matrices)

## ETS, SOLVER (3/4)

## Ion density computation

5 ion species ( $\mathrm{n}_{\text {main }}, \mathrm{n}_{\text {min1 }}, \mathrm{n}_{\text {min2 }}, \mathrm{n}_{\text {imp1 }}, \mathrm{n}_{\text {imp2 } 2}$, index 1,5 ), fully ionized, are generated and deduced from the following equations

## 5 equations

$$
Z_{e f f}=\frac{\sum_{i=1,5} Z_{i}^{2} n_{i}}{n_{e}}, n_{e}=\sum_{i=1,5} Z_{i} n_{i}, \quad \begin{aligned}
& \frac{\mathbf{n}_{\text {min } 1}}{\mathbf{n}_{\text {main }}}=\mathbf{r} 1 \\
& \frac{\mathbf{n}_{\text {min } 2}}{\mathbf{n}_{\text {main }}}=\mathbf{r} 2 \\
& \frac{\mathbf{n}_{\text {imp1 }}}{\mathbf{n}_{\text {imp2 } 2}}=\mathbf{r} 3
\end{aligned}
$$

ETS, SOLVER (4/4)

Source terms


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## V\&V

In collaboration with ISM, comparison with existing codes (ASTRA, CRONOS, JETTO, ...) on JET Shot \#77922

Prescribed quantities

- Heating source (gaussian profile)
- electron density
- equilibrium

Computation

- Ohmic Power, Equipartition, Bremsstrahlung
- Transport Model (NCLASS + B/gB)
- $\Psi, \mathbf{T e}, \mathrm{Ti}$

20 s simulation (1 day of CPU) of the flat-top phase
Good agreement between ETS and existing codes


## II. KEPLER WORKFLOW

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## ETS WORKFLOW

- The KEPLER workflow is adapted for the development
- Reading input data and prescribed profiles from a Cronos input file
- Debug mode allowed: actors are compiled in debug mode, and can be executed under Totalview
- Configure the workflow from Kepler: access to workflow/actors parameters, choice of actors (equilibrium, transport terms, ...)
- Configure a run can be done using ISE but we still are in development phase that's why we don't show the ISE interface
- We tried to limit the workflow complexity by limiting the number of workflow level to 4
- All the configurable workflow parameters are at the toplevel workflow

Schematic view of the workflow


External sources and equilibrium are outside the transport convergence loop
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## EUROPEAN TRANSPORT SOLVER



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## ACTORS

Actors and workflows tested, Actors and workflows under validation

- Equilibrium (HELENA21, HELENA, CHEASE, EMEQ)
- Neoclassic (NCLASS, NEOWES)
- Transport (B/GB, GLF23, COPPITANG, ETAIGB)
- IMP5 workflow (NEMO, RISK)
- NTM
- Sawteeth

The workflow can deal with the 2 ETS solvers (OK in 4.08b but not yet in 4.09a)

All the configurable workflow parameters are at the toplevel

Time parameters

- tbegin_in: 48.1
- tend_in: 50.0
- dtmin_in: $1.0 \mathrm{e}-05$
- dtmax_in: 0.01


## Convergence parameters

- iterationmax_in: 15
- tolerance_in: $1.0 \mathrm{e}-6$
output shot
© runwork_in: 2
- runout_in: runwork_in


## Equation parameters

```
    - ElectronHeatEquation_in: 1
    - IonHeatEquation_in: 1
    - ElectronDensityEquation_in: 0
```


## saving parameters

- sourcecountLH_in: 50
- savenumber_in: 50
- equicount in: 1


## Sources parameters

- sourceLHwithfeedback_in: 0
- PrescribedSourceElectronDensity_in: 1
- PrescribedElectronHeatProfile_in: 1
- PrescribedionHeatProfile_in: 1

Equilibrium parameters

- PrescribeEquilibrium_in: 0
- EquilibriumConvergenceTolerance_in: 5e-2

More sophisticated rules for triggering actors will be implemented and graphical interface will be developed in ISE for configuring these rules.

Change the equilibrium code by double clicking on the equilibrium composite actor


Change the transport coefficients code and neoclassical terms code by double clicking on transport convergence loop


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Change the specific code parameter of an actor by double clicking on the actor and press Edit Code Parameters, for instance the initdata actor (in Initialization subworkflow)


InputFileLocation is the location of the input data file from Cronos. You will see in the pratice part different files examples.

Specific code parameters for the ETS actor (Transport Convergence Loop/TransportSolver/etssolver)


- sca_f is a real scalar between 0.0 (fully implicit) and 0.5 (CrankNicholson schema)
- Persistent memory : compute again the tn matrices or use the persistent memory
- Bremmsthralung : compute or not the internal Bremmsthralung source
- compute_ion_density : compute or not the ion densities

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## HOW TO ADD YOUR ACTOR 1/4

Adding an actor which is not an equilibrium code nor a transport coefficient nor a neoclassical code is relatively complex and it is better to discuss first with the ETS developpement team

The Kepler bundle is a record that contains the references to the plasma state that are needed for your code. This bundle is transmitted from one composite actor to another. It is currently composed by the following but will evolve.
time composed by cvg composed by ual composed by
: tn,tnp1,dt,iterloop
: eps,iter,count,bestvalue
: coreptn,coreptnp1,neotn,neotnp1, corettn,corettnp1,equitn,corestn,antentn prescribedual composed by : equi,corep,coret,anten

We use different occurrences: Occurrence $2=$ tn
Occurrence 3 = tnp1
Occurrence 1 = prescribed cpo
Occurrence 0 = final results

## HOW TO ADD YOUR ACTOR 2/4

For an equilibrium actor
-Go to equilibrium/equilibrium convergence/compute new equilibrium subworkflow

The output occurence must be 3

-Replace one equilibrium actor by your actor and don't forget to set the output equilibrium occurrence to 3

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## HOW TO ADD YOUR ACTOR 3/4

For a neoclassical actor
-Go to Transport Convergence Loop/Neoclassical Terms subworkflow

-Replace one neoclassical actor by your actor and don't forget to set the output neoclassical occurrence to 3

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## HOW TO ADD YOUR ACTOR 4/4

For transport coefficients actor
-Go to Transport Convergence Loop/Transport Coefficients subworkflow


The output occurence must be 3
-Replace one transport coefficents actor by your actor and don't forget to set the output neoclassical occurrence to 3

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## An example of inserting NTM actor in the ETS workflow

Work in progress done with S.NOWAK and O.SAUTER


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# III. Visualization and Post-treatement 

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You will use in the practice different methods to visualise results stored in the UAL database :

- freely:
- using matlab interpretor
- using python interpretor
- among predefined choices :
- using VCOS [C-program]

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## Part 1 : UAL-access

## INTEGRATED TOKAMAK MODELLING

```
Python syntax
from vmod import *
shot = 10
run =4
ob = ual.itm(shot,run,shot,0)
ob.open()
ob.coreprofArray.get()
ob.close()
nt = len(ob.coreprofArray.array)
nrho = len(ob.coreprofArray.array[0].te.value)
# comment line must begin with a < # » character
# array indices start from 0 and are inclosed in square brakets
time = np.zeros([nt])
ip = np.zeros([nt])
| = np.zeros([nt])
psi = np.zeros([nt,nrho])
rho = np.zeros([nt,nrho])
te = np.zeros([nt,nrho])
jtot = np.zeros([nt,nrho])
legs = []
for kt in range(nt) :
    coreprof = ob.coreprofArray.array[kt]
    time[kt] = coreprof.time
    psi[kt,:] = coreprof.psi.value[:]
    rho[kt,:] = coreprof.rho tor_norm[:]
    ip[kt], = coreprof.globālpāram.current_tot
    li[kt] = coreprof.globalparam.li
    te[kt,:] = coreprof.te.value[:]
    jtot[k't,:] = coreprof.profilesld.jtot.value[:]
    legs.append("t = %(#).3f s" % {'#': time[kt]})
# you must indent (using tab) loop body
```


## Matlab syntax

```
shot = 10;
run = 4;
ob = euitm open('euitm',shot,run);
acoreprof = euitm get(ob,'coreprof');
euitm_close(ob)
nt = length(acoreprof);
nrho = length(acoreprof(1).te.value);
```

$\%$ comment line must begin with a « \% » character
\% array indices start from 1 and are inclosed in parenthesis
\% array initialisation are not mandatory
legs $=\{ \}$;
for $k t=[1: n t]$
coreprof $=$ acoreprof(kt);
time (kt)
psi(kt,:) = coreprof.psi.value;
rho(kt,:) = coreprof.rho tor_norm;
ip(kt) = coreprof.globālpāram.current_tot;
li(kt) = coreprof.globalparam.li;
te(kt,:) = coreprof.te.value;
jtot(kt,:) = coreprof.profiles1d.jtot.value;
legs $=\left\{\right.$ legs $\{:\}$ ['t $=$ ' num2str(time(kt)) ' $\left.\left.\mathrm{s}^{\prime}\right]\right\}$;
end

## Python syntax

figo $=$ plt.figure(1)
figo.clf(t)
plt.plot(time, ip,'-or')
 pt.title("shot $\%(A) d, r u n \%(B) d:$ plasma current" \%\{"A":shot,"B":run\})) bit: y abel('in') pt:.xabel('time')
fig1 $=$ gltt.figure(2)
plt.plot(nn.transpose(psi),np.transpose(jtot))
bit titlen (hiogs \%(shot)d,run \%(run)d : total current" \% \{"shot":shot,"run":run\})) ti-y plt.xabel("psi')
plt.show()

## Matlab syntax

```
figure(1)
cif
plot(time,ip,'-or')
```



```
ylabe t'in'
y xabel('time')
figure(2)
plot(psi.', jtotot.')
ligend(liegs.)
y apel('jtot')
xabel('psi')
```

vmod.py contain python modules needed to read UAL-database and plot:
\#!/usr/bin/env python
import matplotlib.pyplot as plt \# plot module import numpy as np import ual
\# n-dimensional array module
\# ual database access module

Documentation can be found at url :

- python [current installed version] : http://docs.python.org/release/2.5.1/
- matplotlib : http://matplotlib.sourceforge.net/
- numpy: http://numpy.scipy.org/
- ual :
https://www.efda-itm.eu/ITM/imports/isip/public/isip_UAL_User_Guide.pdf

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VCOS an interactive visualisation compiled program :

- C-language
- UAL access through C++ module
- communicate whith a CGI server (spcgi.cgi) :
- => XHTML-Forms with embedded SVG on Firefox browser
(XML compliant, W3C specifications )
- <= read edited parameters using POST method


## Actual features : Validation of an ETS simulation by comparison with CRONOS

- read UAL-data (coreprof-CPO)
- read data from Matlab files after conversion in ASCII-file :
cronos-output.mat --[zmat2txt.m script]--> cronos-output.dat
- compare data for various shot / run or CRONOS output Matlab files :
- profiles at final time for 1D data
- time evolution for global data

To be done :

- save SVG content on file for further processing (Inskape, Gimp ...)
- implement a time navigator

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$$

## Launching VCOS: vcos [-file cronos-output.dat]



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## IV. Practice

## GET ETS workflow and actors

Get the material for the ETS workflow by copying GARCHING2011 on your HOME directory or WORK directory
cp -r ~huynh/public/GARCHING2011 ~

In this directory, you can find :

- the instructions of importing ETS actors are written in the README file.
- the ETS workflow stored in WORKFLOWS_AND_ACTORS
- different input files generated and stored in INPUT_FILES


## EXERCISE 1 <br> (current diffusion equation, interpretative mode)

Modify the workflow:

- Update the InputFileLocation parameter of the "initdata" actor with the location of file_ETS_77922
- Change the tend_in value to 48.2 s
- Put the interpretative mode (solved only $\Psi$ )
- Verify the mode "Prescribed external source"
- Modify the runwork_in to 3
- save the workflow into EXE1

Execute the workflow

- Don't forget to type itmgo befor launching the execution

Visualization of the results (comparison with CRONOS) :

- vcos -file ISM_77922_new_9_ter_resultat.dat

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# EXERCISE 2 <br> (more diffusion equations, predictive mode) 

Modify the workflow :

- Put the predictive mode (solved $\Psi, \mathrm{Te}, \mathrm{Ti}$ )
- Verify the transport model (Bohm/gyroBohm)
- modify the runwork_in to 4
- save the workflow into EXE2

Execute the workflow

Visualization of the results (comparison with CRONOS and with the interpretative mode)

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## EXERCICE 3 <br> (predictive mode, equilibrium convergence)

Modify the workflow :
-Keep the predictive mode (solved $\Psi, \mathrm{Te}, \mathrm{Ti}$ )
-Decrease the dtmin_in to 0.03
-Computation of the equilibrium (put equilibrium time step to 1)
-Verify prescribed equilibrium
-Modify the runwork_in to 5
-Save the workflow into EXE3

Execute the workflow

Visualization of the results (comparison with CRONOS and with the interpretative mode)

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## EXERCISE 4 (predictive mode, source term + feedback)

Modify the workflow :

- keep the predictive mode (solved $\Psi, \mathrm{Te}, \mathrm{Ti}$ )
- prescribed equilibrium
- turn on LH source module, allowed feedback control
- modify the runwork_in to 6
- save the workflow into EXE4

Execute the workflow

Visualization of the results (comparison with CRONOS and with the interpretative mode)

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